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Charles Currat LBNL

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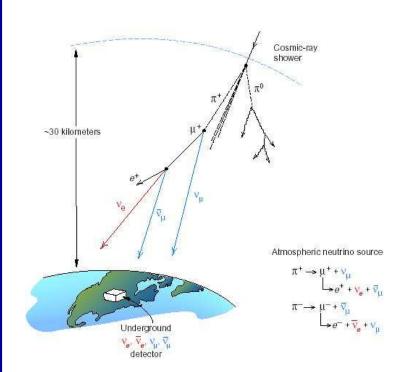
Atmospheric neutrinos studies in SNO

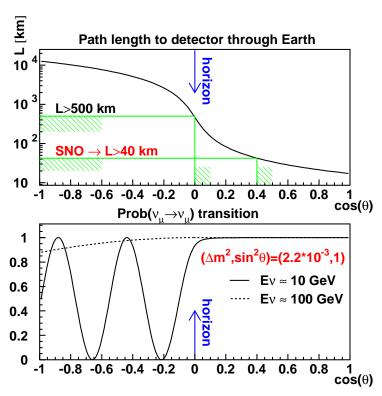


Atmospheric neutrinos in SNO



Challenging for SNO since volume is much smaller than SuperK: 2.7 kton vs 22 kton fiducial (50 kton total). But at its depth SNO is in a unique position amongst underground detectors.





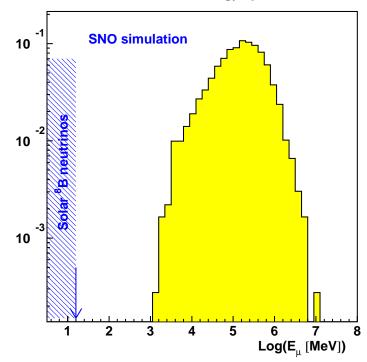


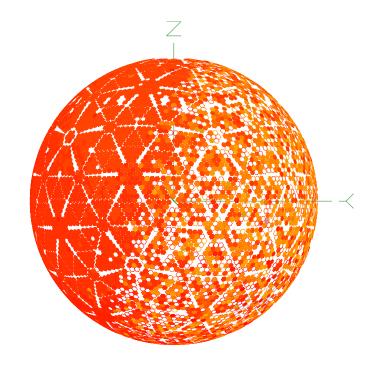
Muons in SNO ____



We want to measure through-going muons ($E_{\mu}>10$ GeV). 3D-track reconstruction uses PMTs timing.

Cosmic muons energy spectrum





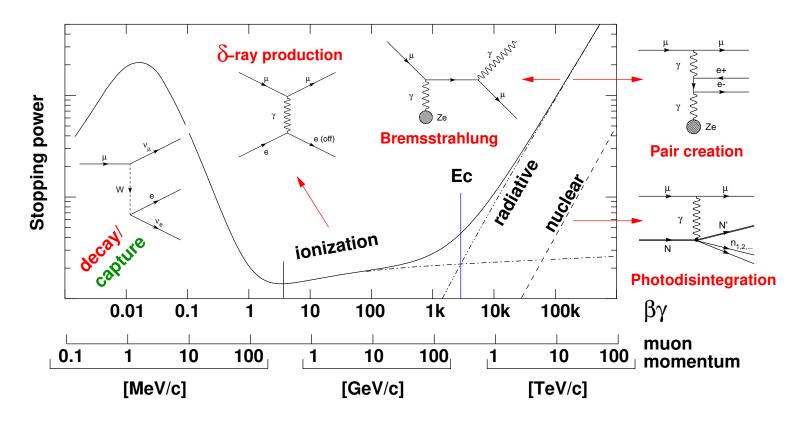
Run: 1 GTID: 33



Software side



Important effort on muons simulation in SNO Monte Carlo.



From O(10 TeV) down to explicit thermalization of spallation products (neutron 1/40 eV) the same data structure accommodates 15 orders of magnitude in energy!!

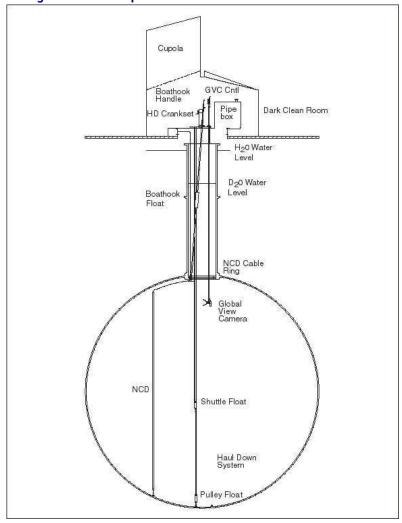


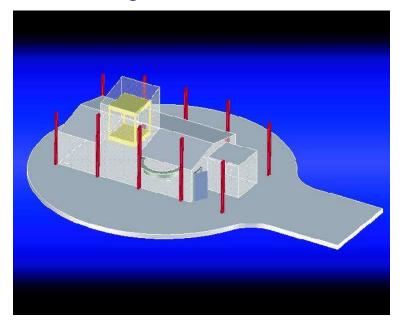
Hardware side _____

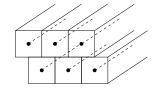


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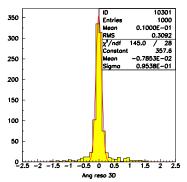
Project: independent calibration of muons with tracking chambers.







Cell size: 7.5 cm x 7.5 cm Longitudinal resolution: 5 mm Transverse resolution: 500 microns Stations: 2, 2 layers each, 2 m apart





Chambers



We found chambers used for HEP test beams (FNAL/IUCF/JLab). Some engineering work required for SNO needs (scintillator pads, support structure).



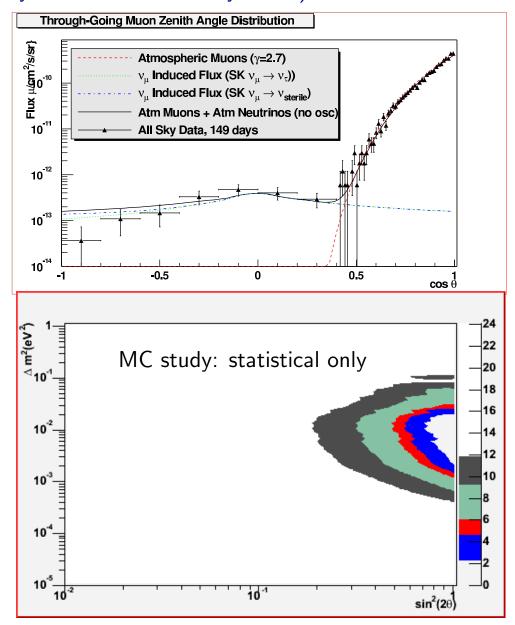




Atmospheric neutrinos in SNO



Preliminary analysis with 150 days of data (courtesy of N. Tagg) Perspective at SNO with 730 days of data (probably over 800 days available, ultimately $\sim \times 2$)



stop/thru analysis à la SuperK under investigation (bin over horizon)